

-	OGLE16aaa	01:07:20.88	-64:16:20.70	TDE candidate in a weak AGN	Wyrzykowski et al. (2017)
AT 2019eve	ZTF19aatylnl/Gaia19bti	11:28:49.650	+15:40:22.30	Persistent H_{α} line one year post peak.	van Velzen et al. (2021) ,
AT 2021lwx	ZTF20abrbeie	21:13:48.405	+27:25:50.46	Excluded from the Hammerstein et al. (2023) as an outlier.	Hammerstein et al. (2023)
AT 2024kmq	ZTF24aapvieu/ATLAS24jwz/GOTO24dqz	12:02:37.273	+35:23:35.22	Ultraluminous, long duration transient	Subrayan et al. (2023)
-	Swift J1112+82	11:11:47.32	-82:38:44.20	Luminous, fast, red transient	Ho et al. (2025)
				Likely jetted TDE	Brown et al. (2015)
AT 2019fdr	ZTF19aatubsj/Gaia19bsz/ ATLAS19lkd/PS19dar	17:09:06.859	+26:51:20.50	Initially classified as SLSN-II on TNS.	Yan et al. (2019) ,
AT 2019baf	ZTF19aaciohh	17:52:00.146	+65:37:35.99	Later classified as TDE candidate, however	Frederick et al. (2021)
AT 2019cmw	ZTF19aaniqrr/PS19cws/ ATLAS19kel	18:48:39.472	+51:00:48.66	an extreme AGN flare origin is also possible.	
AT 2020abri	ZTF20acvezvs/ATLAS20biig	13:29:17.278	+19:40:15.34	Sect. 2.2	Yao et al. (2023)
				Sect. 2.2	Yao et al. (2023)
AT 2023adr	ZTF22abzajwl/ Gaia23aui/ATLAS23bcm	14:36:19.830	+32:23:16.48	Initially classified as a SLSN candidate. Later	Perley et al. (2023) ,
				it was proposed that maybe this is a TDE due to	Shlentsova et al. (2024) ,
-	PS1-11af	09:57:26.82	+03:14:01.00	its blue color. Recently identified as a	Llamas Lanza et al. (2024)
-	ASASSN-14ko	05:25:18.13	-46:00:20.34	repeated pTDE candidate	Chornock et al. (2014)
-	PTF-09axc	14:53:13.06	+22:14:32.2	Optical/UV TDE candidate, could be TDE	Payne et al. (2021)
				Repeating partial TDE is the most probable,	
				but not the only one.	Arcavi et al. (2014)
SN 2016ezh	PS16dtm	01:58:04.739	-00:52:21.74	Weak AGN activity could not be excluded.	Tomasella (2016)
				Classified as SN II on TNS.	
				Later introduced as candidate TDE	
				in a Seyfert 1 host galaxy.	Blanchard et al. (2017)
-	ASASSN-15lh	22:02:15.39	-61:39:34.60	Ambiguous transient, displaying features	Dong et al. (2016) ,
				consistent with both SLSN and TDEs.	Leloudas et al. (2016) ,
-	DES14C1kia	03:34:47.49	-26:19:35.0	Likely TDE candidate, with no X-ray or	Godoy-Rivera et al. (2017)
				radio emission.	Foley et al. (2015) ,
					Yu et al. (2015) ,
					Ravi & Shannon (2015)
-	D3-13	14:19:29.81	+52:52:06.37	Optical/UV TDE candidate, however the light	Gezari et al. (2006) ,
-	PTF10iya	14:38:41.00	+37:39:33.6	curve is incomplete and the flare's position	Gezari et al. (2008)
-	SDSS J1323	13:23:41.97	+48:27:01.26	is off-center slightly from the host galaxy	
-	SDSS J1342	13:42:44.42	+05:30:56.14	An unusual AGN scenario cannot be ruled out.	Cenko et al. (2012)
				Potential TDE candidate, however	Esquej et al. (2007) ,
				LLAGN interpretation cannot be fully excluded	Esquej et al. (2008)
				TDE candidate based on coronal emission line	Yang et al. (2013) ,
				detection	Dou et al. (2016)
-	SDSS J0952/ 2MASSJ09520955+2143132	09:52:09.56	+21:43:13.24	Possible TDE candidate, however this source	Komossa et al. (2008) ,
				could host a permanent LLAGN.	Palaversa et al. (2016) ,
-	PS1-10adi	20:42:44.749	+15:30:32.24		Dou et al. (2016)
-	OGLE17aaj	01:56:24.93	-71:04:15.7	Candidate TDE in an AGN	Jiang et al. (2019)
				Potential TDE candidate, however it could	
				represent a change in AGN accretion flow.	Gromadzki et al. (2019)
AT 2019aalc	ZTF19aaejtoy	15:24:16.664	+04:51:19.05	Initially classified as TDE, however its	Veres et al. (2023) ,
				properties can be explained by a typical Seyfert 1.	Guolo & Gezari (2023)

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